

The IXO WFI, HXI and HTRS



Overview over the WFI and HXI
concepts and activities

Lothar Strüder, MPE, HLL

Outline:

WFI

- The DePFET active pixel sensor
- The WFI layout and properties
- The WFI status and technology development plan

HXI

- The double sided CdTe and DSSD hard X-ray Imager
- The HXI concept and the expected properties

HTRS

- The high speed SDD
- The expected properties

with input
from

WFI (Lothar Strüder)

- **MPI-HLL (MPE and PNSensor)**
- **Universties of Tübingen, Darmstadt and Nürnberg-Erlangen**
- **University of Leicester**
- **Politecnico di Milano**
- **University of Osaka**
- **MIT, CfA and Penn State University**

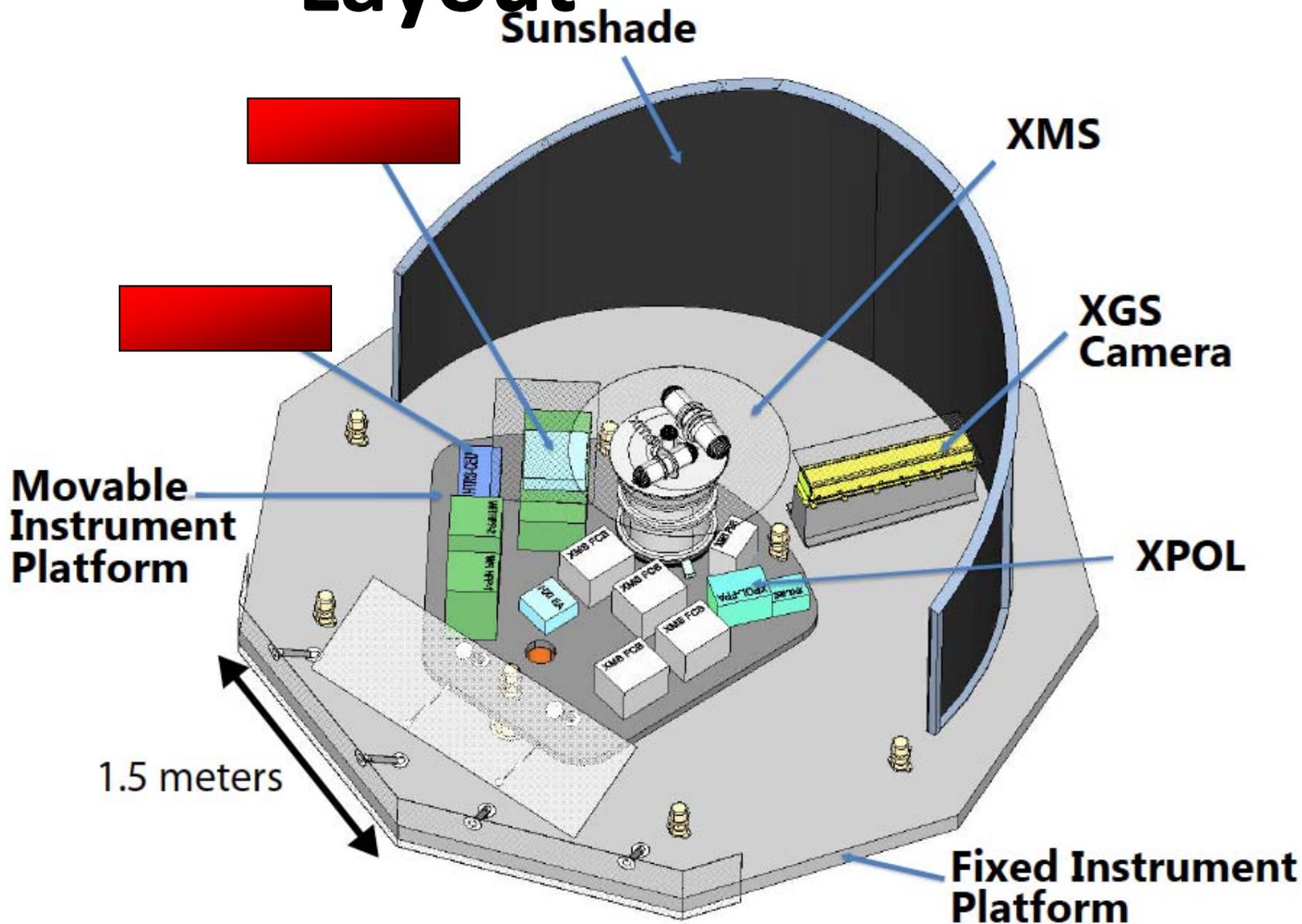
HXI (Tad Takahashi)

- **JAXA, ISAS**
- **CEA, CNES**
- **University of Tokio**

HTRS (Didier Barret)

CESR, CNES, MPE/PNS
University of Tübingen,
University de Genève
Univ. of Nürnberg-Erlangen
Politecnico di Milano

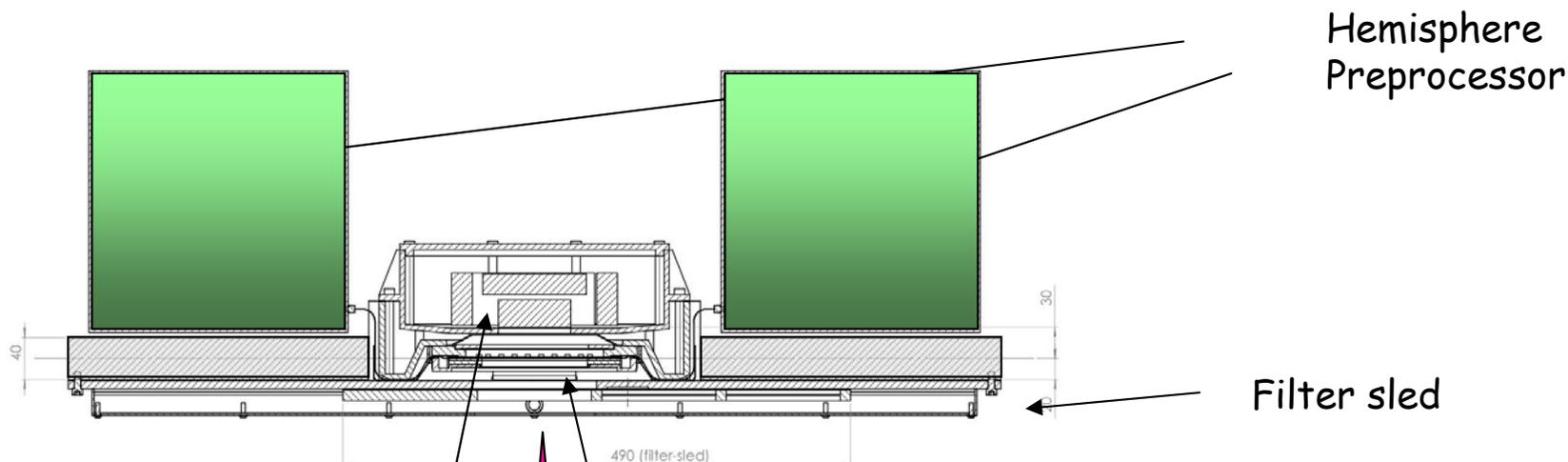
Focal Plane Layout



Mirror Effective Area	<p>3 m² @1.25 keV</p> <p>0.65 m² @ 6 keV with a goal of 1 m²</p> <p>150 cm² @ 30 keV with a goal of 350 cm²</p>	<p>Black hole evolution, large scale structure, cosmic feedback,</p> <p>Strong gravity, EOS</p> <p>Cosmic acceleration, strong gravity</p>
Spectral Resolution (FWHM)	<p>$\Delta E = 2.5$ eV within 2 x 2 arc min (0.3 – 7 keV)</p> <p>$\Delta E = 10$ eV within 5 x 5 arc min (0.3 – 7 keV)</p> <p>$\Delta E = 150$ eV at 6 keV within 18 arc min diameter (0.1 - 15 keV)</p> <p>$E/\Delta E = 3000$ (0.3–1 keV) with an area of 1,000 cm² and a goal of 3000 cm² for point sources</p> <p>$\Delta E = 1$ keV within 8 x 8 arc min (10 – 40 keV)</p>	<p>Black Hole evolution,</p> <p>Large scale structure</p> <p>Missing baryons using tens of background AGN</p>
Angular Resolution	<p>≤ 5 arc sec HPD (0.1 – 7 keV)</p> <p>30 arc sec HPD (7 - 40 keV); goal of 5 arc sec</p>	<p>Large scale structure, cosmic feedback,</p> <p>black hole evolution, missing baryons</p>
Count Rate	<p>1 Crab with >90% throughput. $\Delta E < 150$ eV @ 6 keV (0.1 – 15 keV)</p>	<p>Strong gravity, EOS</p>
Polarimetry	<p>1% MDP on 1 mCrab, 100 ksec, 3σ, 2 - 6 keV</p>	<p>AGN geometry, strong gravity</p>
Astrometry	<p>1 arcsec at 3σ confidence</p>	<p>Black hole evolution</p>
Absolute Timing	<p>50 μ sec</p>	<p>Neutron star studies</p>

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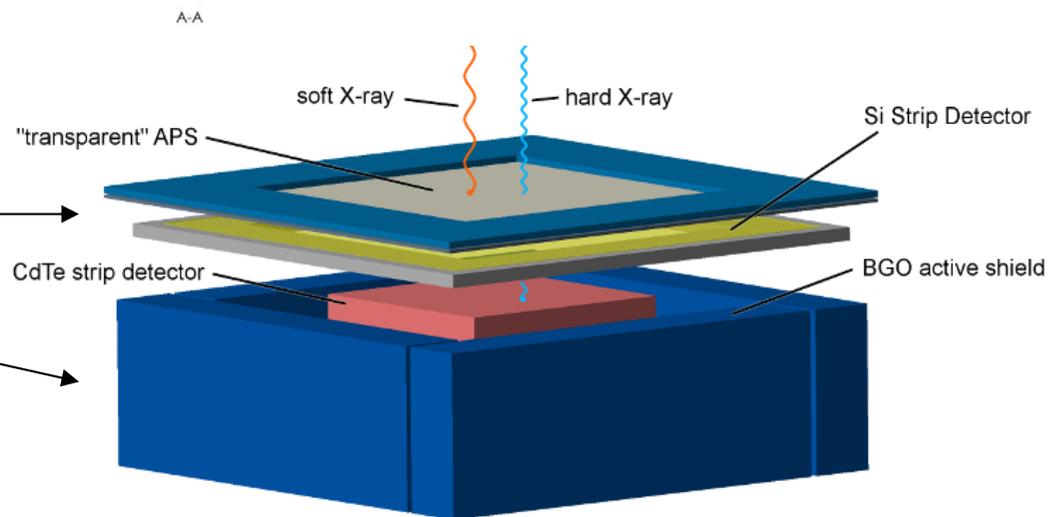
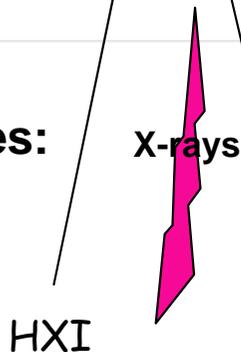
WFI/HXI Mechanical Layout



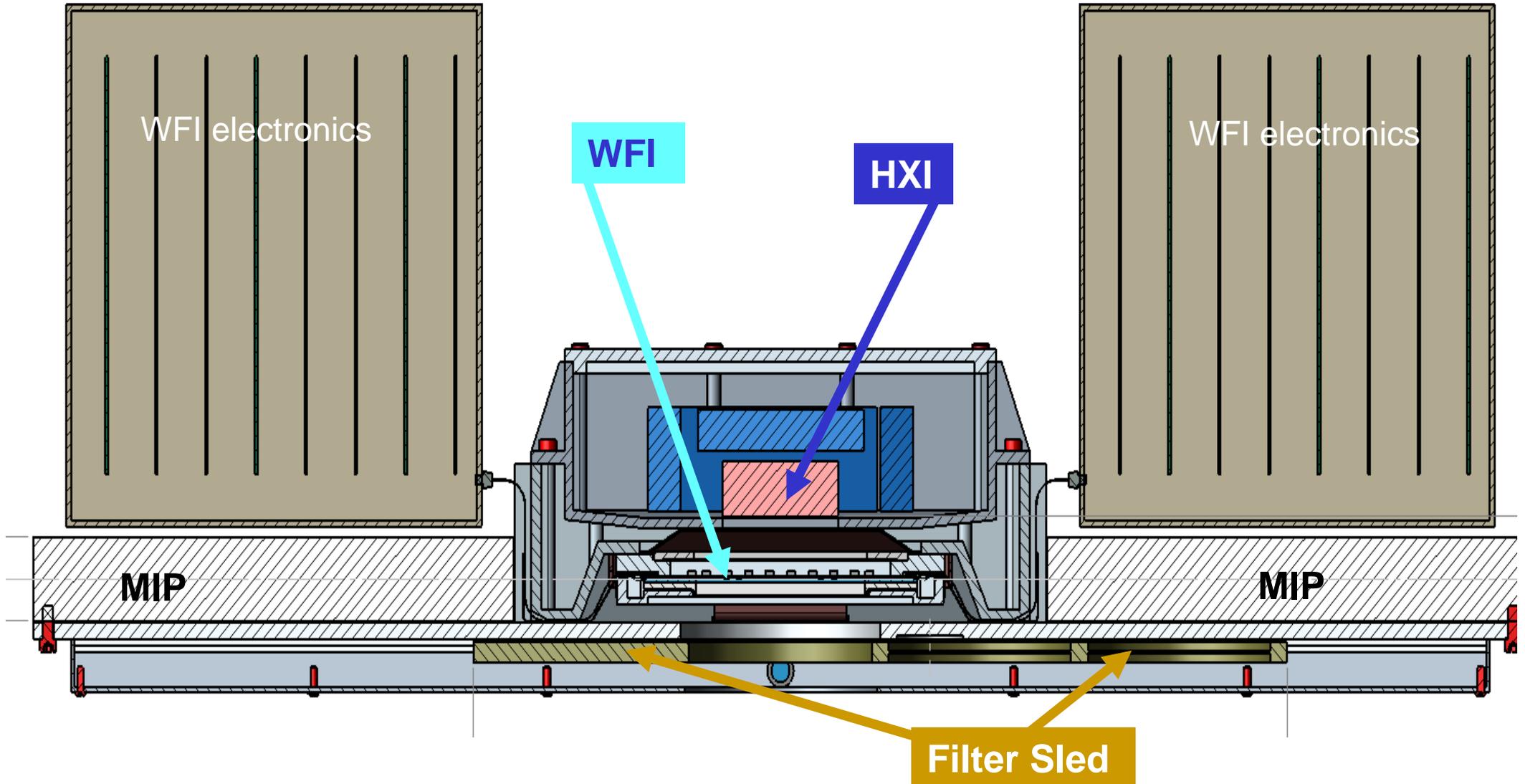
Wide field imagers
for two energy ranges:

WFI: 0.05 – 15 keV
FOV: 18x18 arcmin

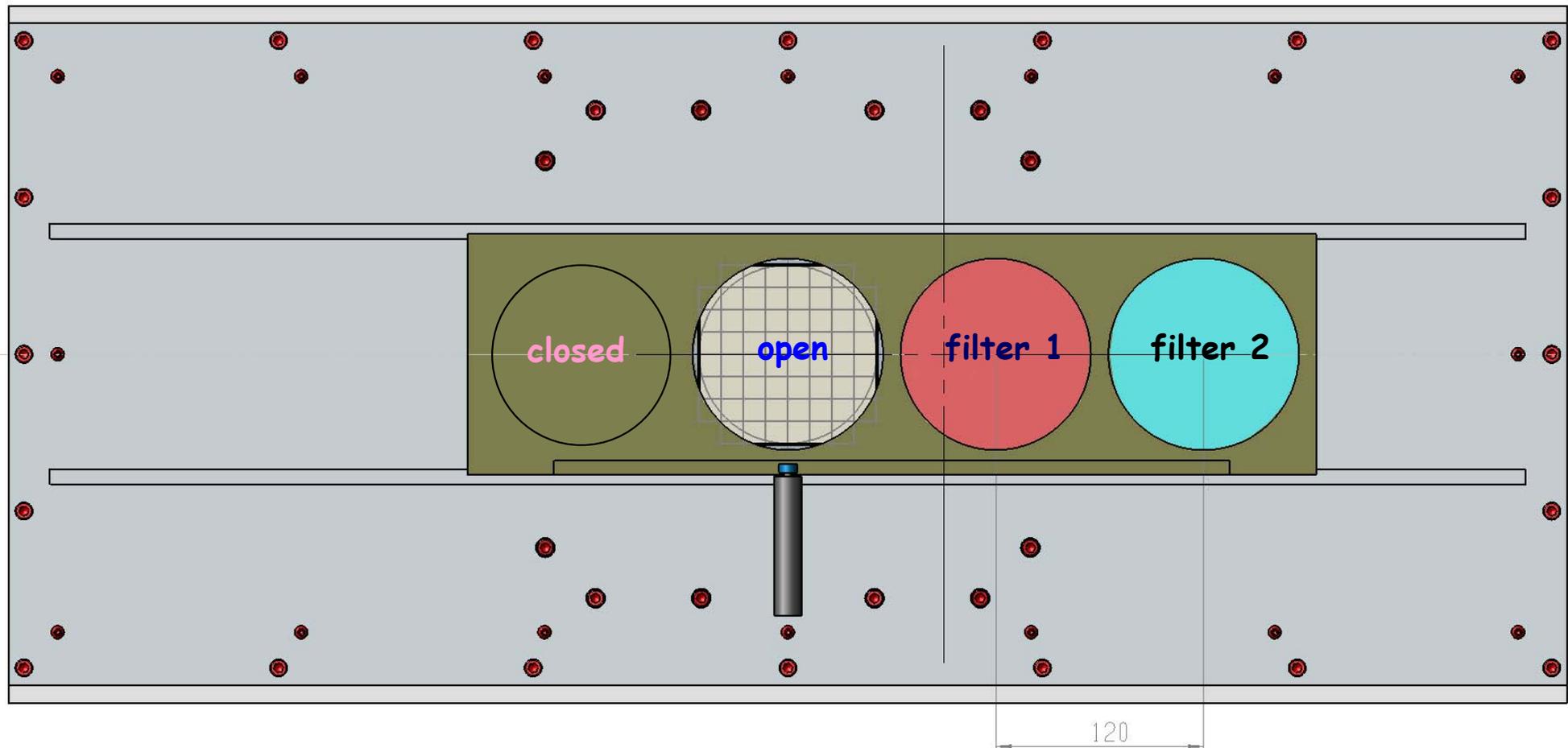
HXI: 5 – 50 keV
FOV: 8x8 arcmin



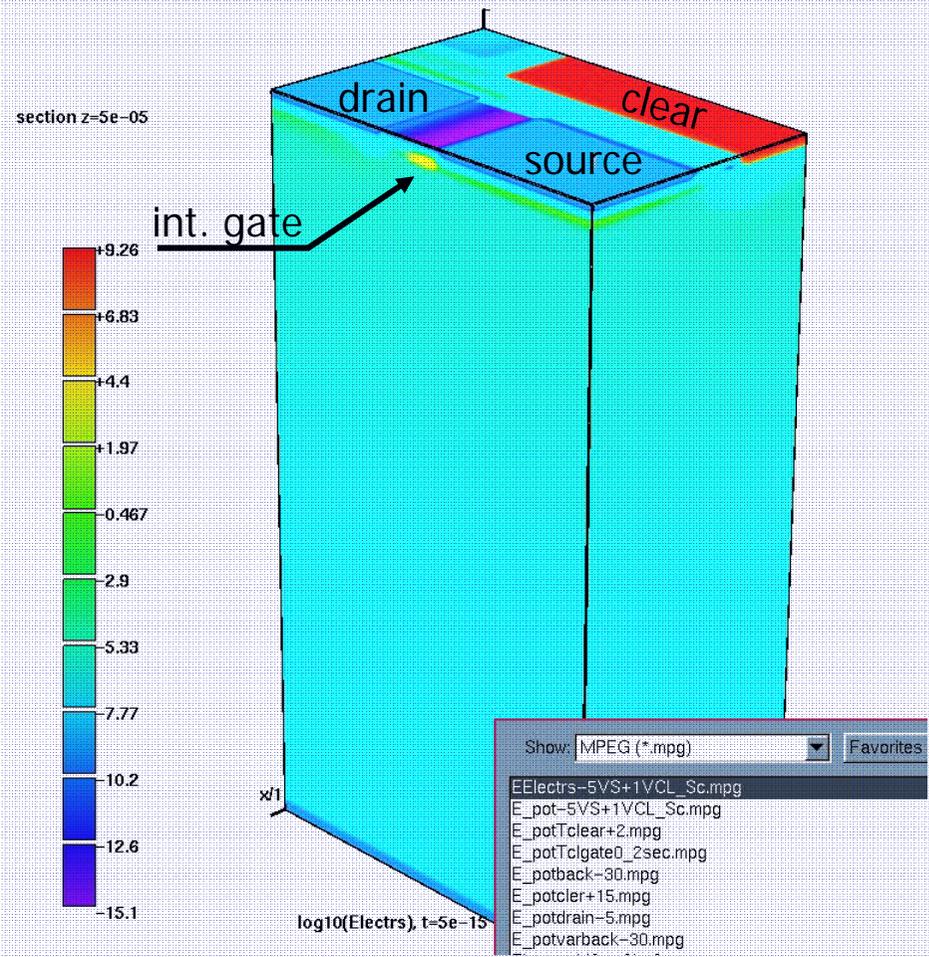
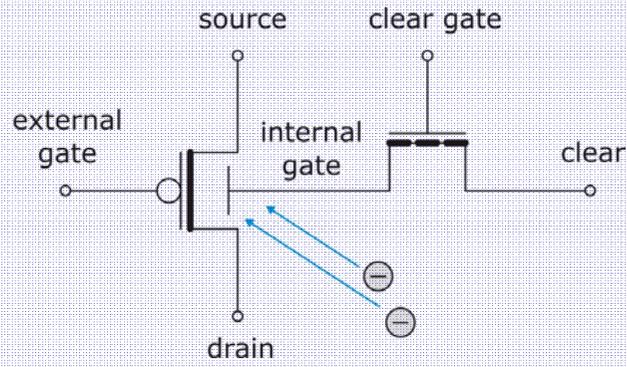
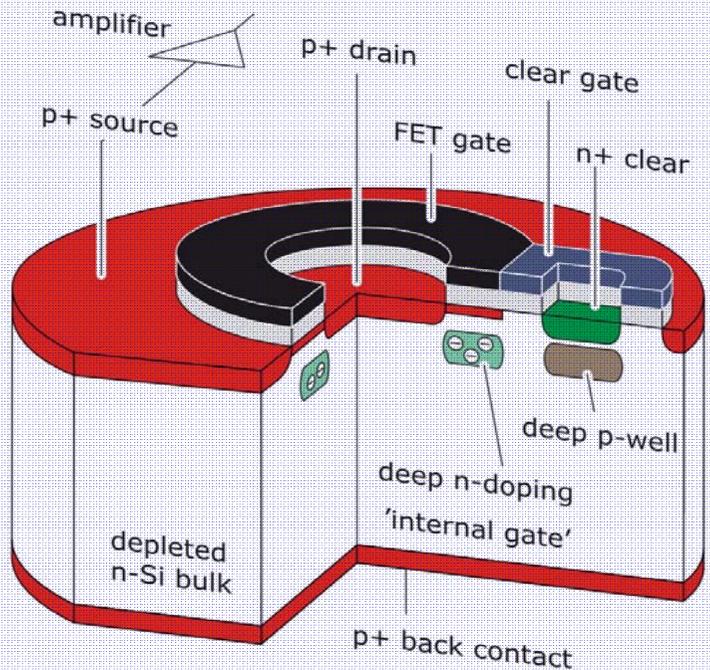
Thermal & Mechanical Concept



Filter Sled

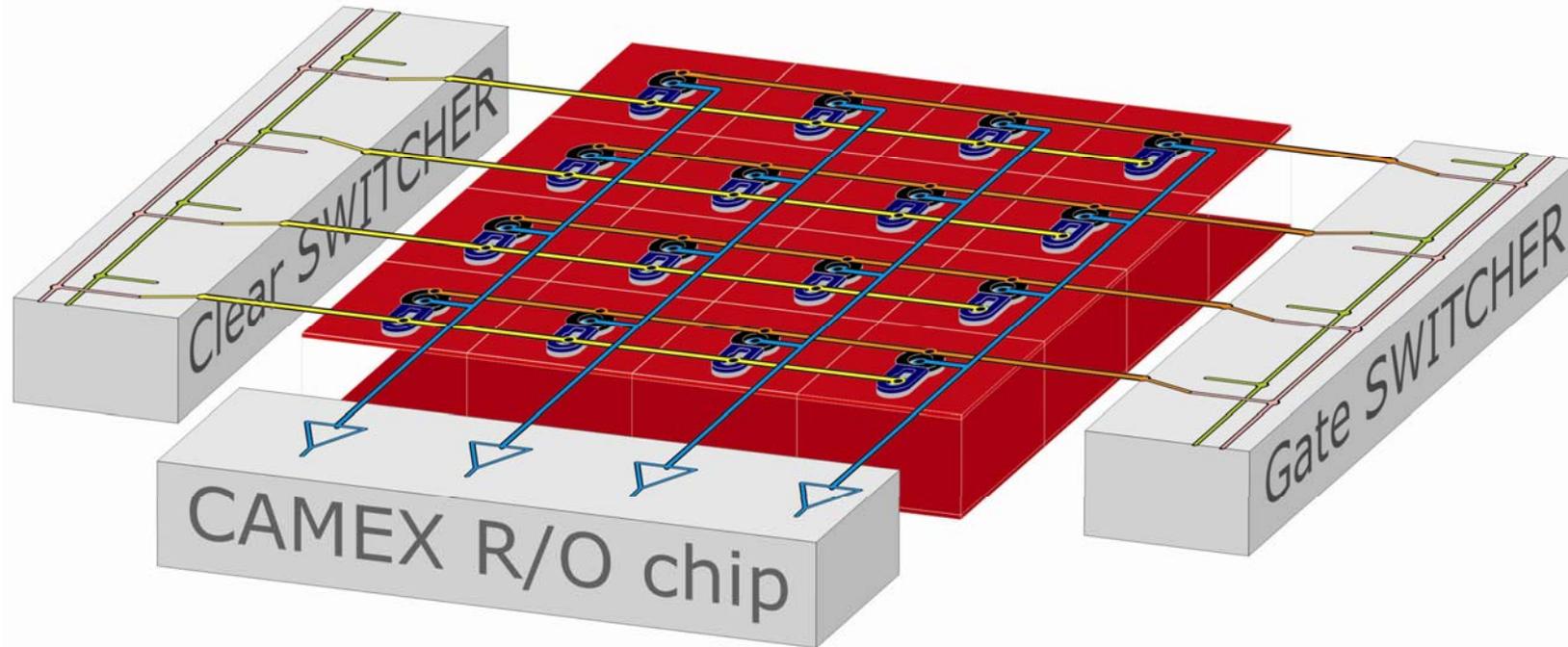


3-dim simulation of the DEPFET



TeSCA 3D Simulation by K.Gärtner, WIAS, Berlin

DEPFET Active Pixel Sensor Array



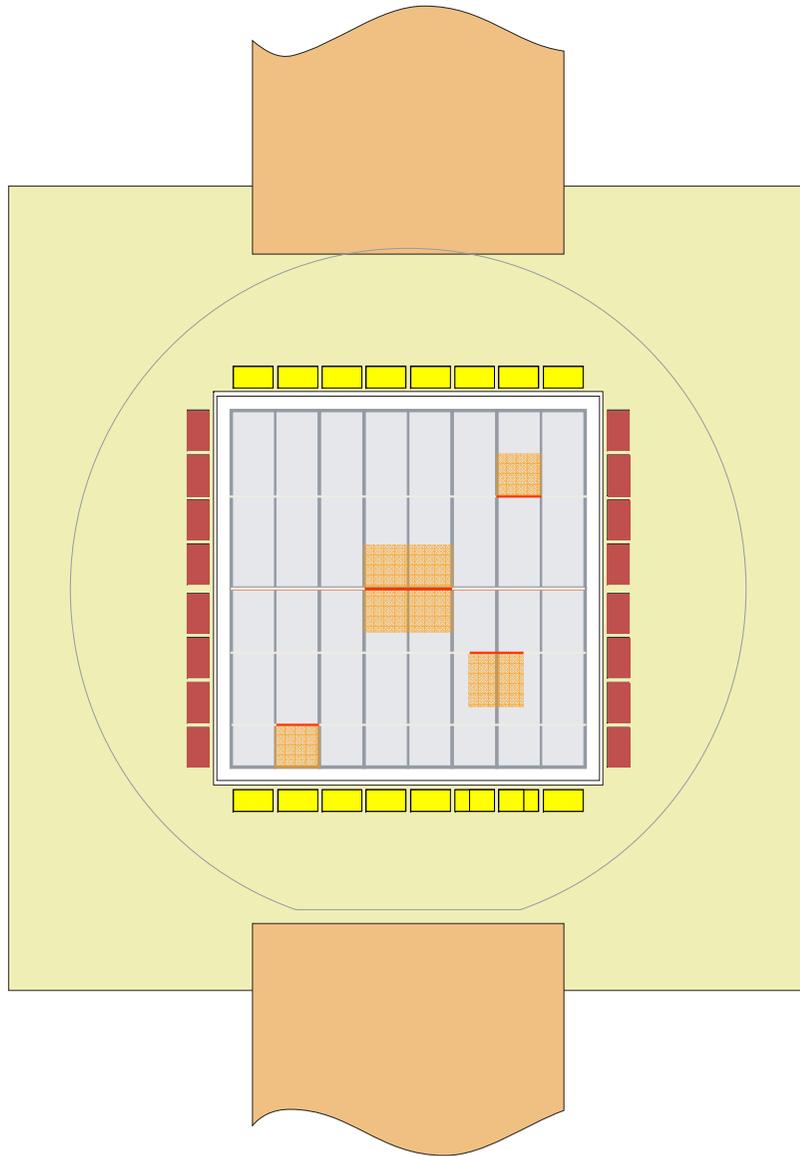
- **matrix organisation**

- common back contact
- » **thin, homogeneous entrance window**
- » **fill factor 100 %**
- row-wise connection of gate, clear, clear gate
- column-wise connection of source / drain
- » **individually addressable pixels**
- » **windowing option**

- **operation philosophy**

- one active row
- all other pixels turned off
- » **low power consumption**
- all operations in a row in parallel
- » **fast processing**

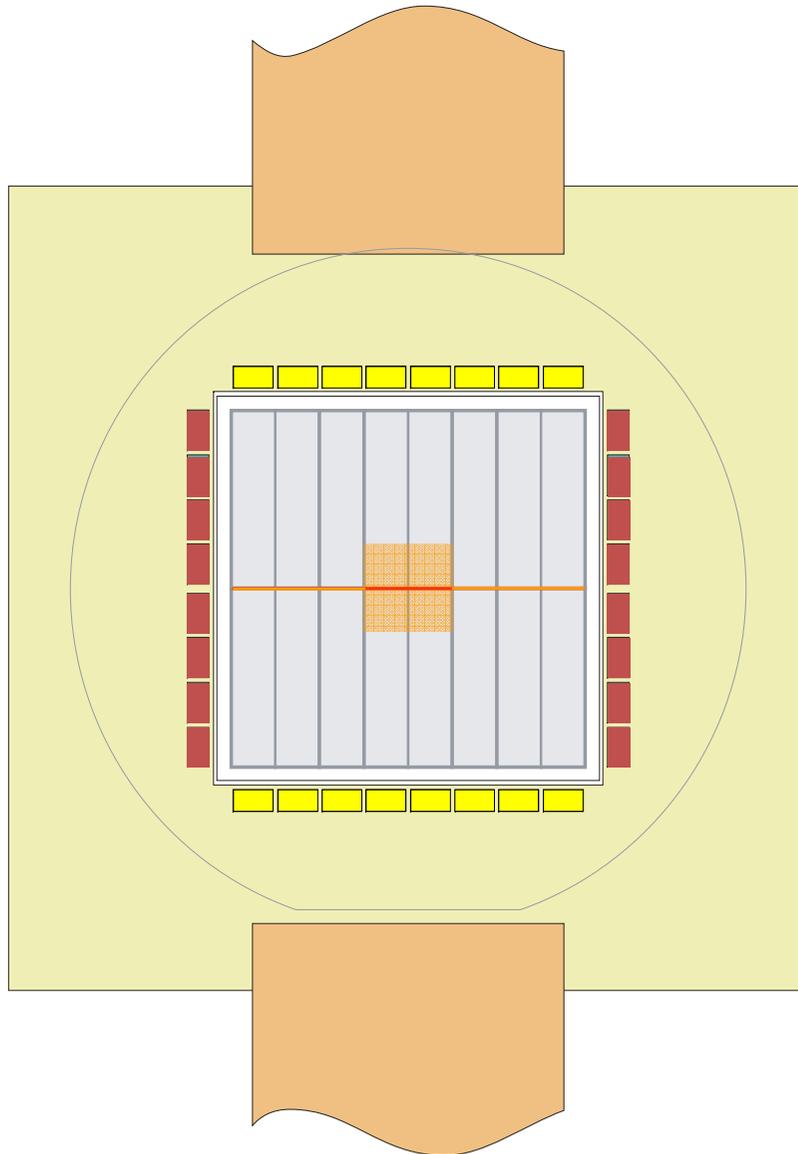
WFI readout concept



■ *Readout modes:*

- **Full frame mode:** Parallel readout of both hemispheres on full width
- **ROI mode:** Define ROI, read out repetitively with high framelet rate
- Information of entire row is acquired, but information from outside ROI is discarded in preprocessing
- Arbitrary position anywhere on the sensor
- Simultaneous readout of disjunkt ROIs on different hemispheres
- With next generation of ICs:
- On-the-fly selection of ROIs / switch between ROIs
- ROIs exceeding sector borders

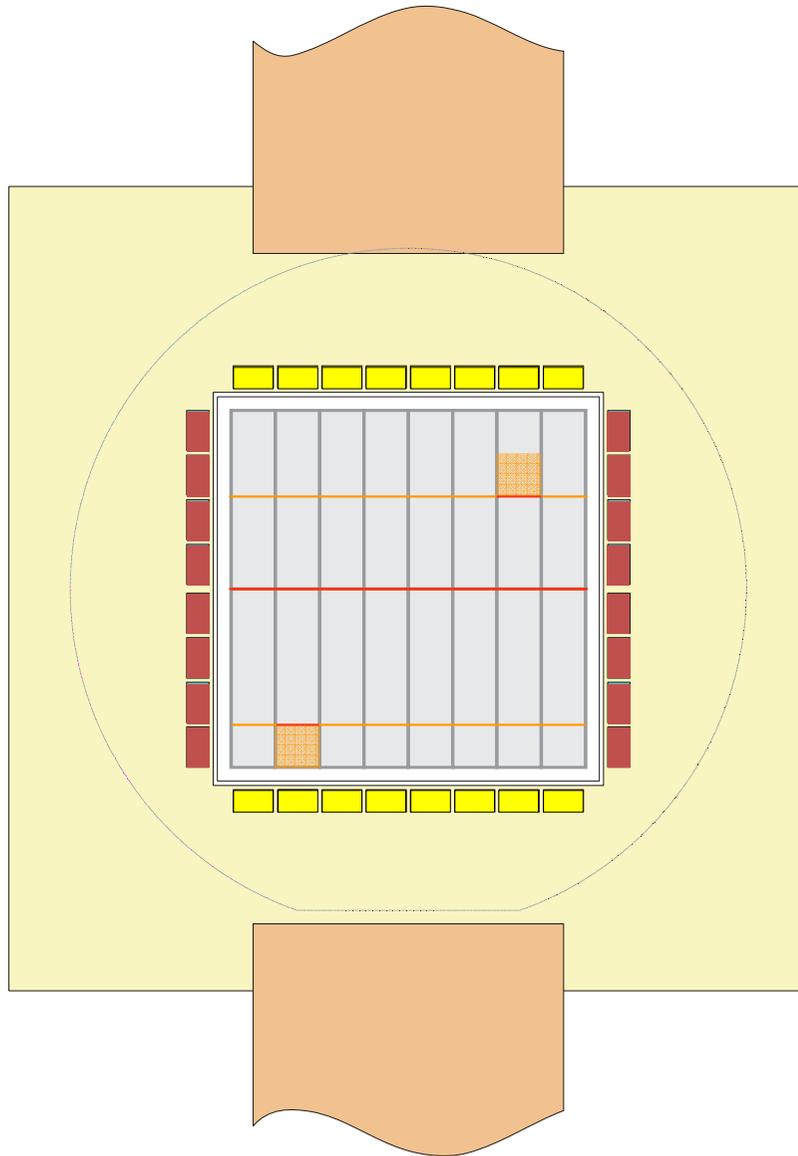
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- **Window mode:** Acquire fully sized window strip (anywhere on FPA) repetitively
- Read rest of frame with reduced framerate

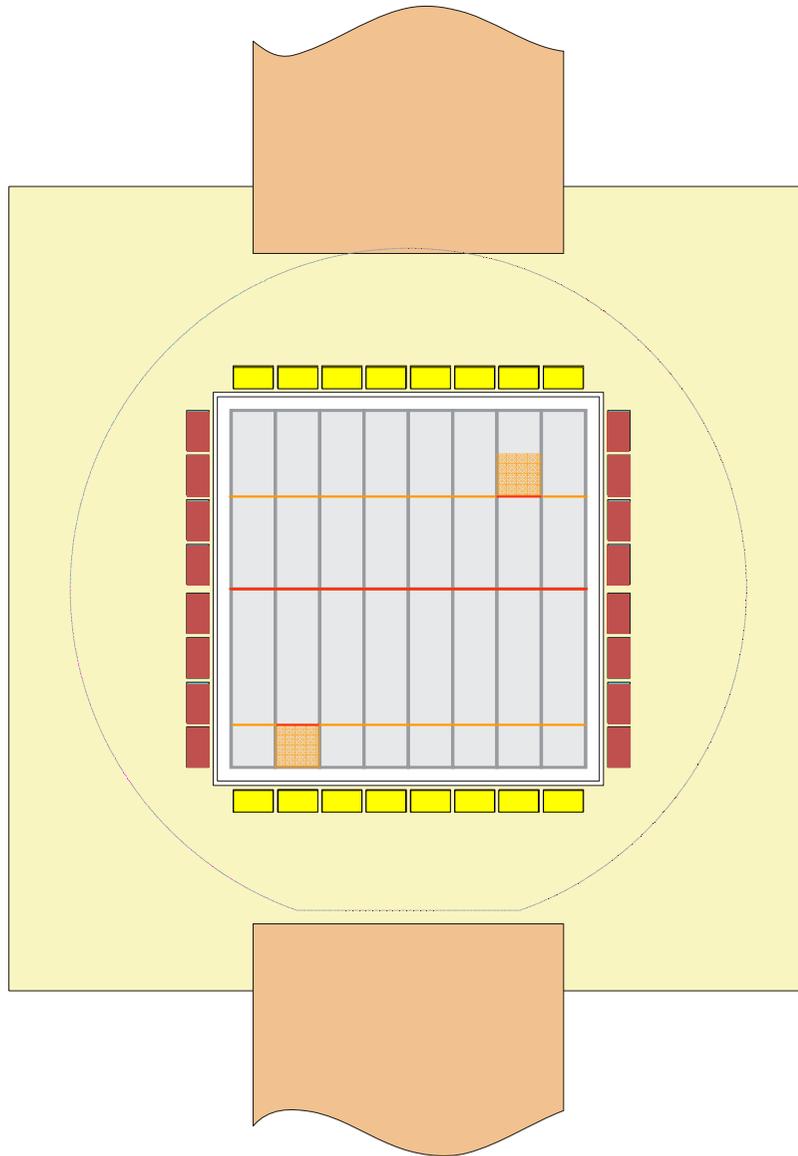
WFI readout concept



■ *Readout modes:*

- **Window mode:** Acquire fully sized window strip (anywhere on FPA) repetitively
- Read rest of frame with reduced framerate
- Different ROIs on arbitrary positions on FPA

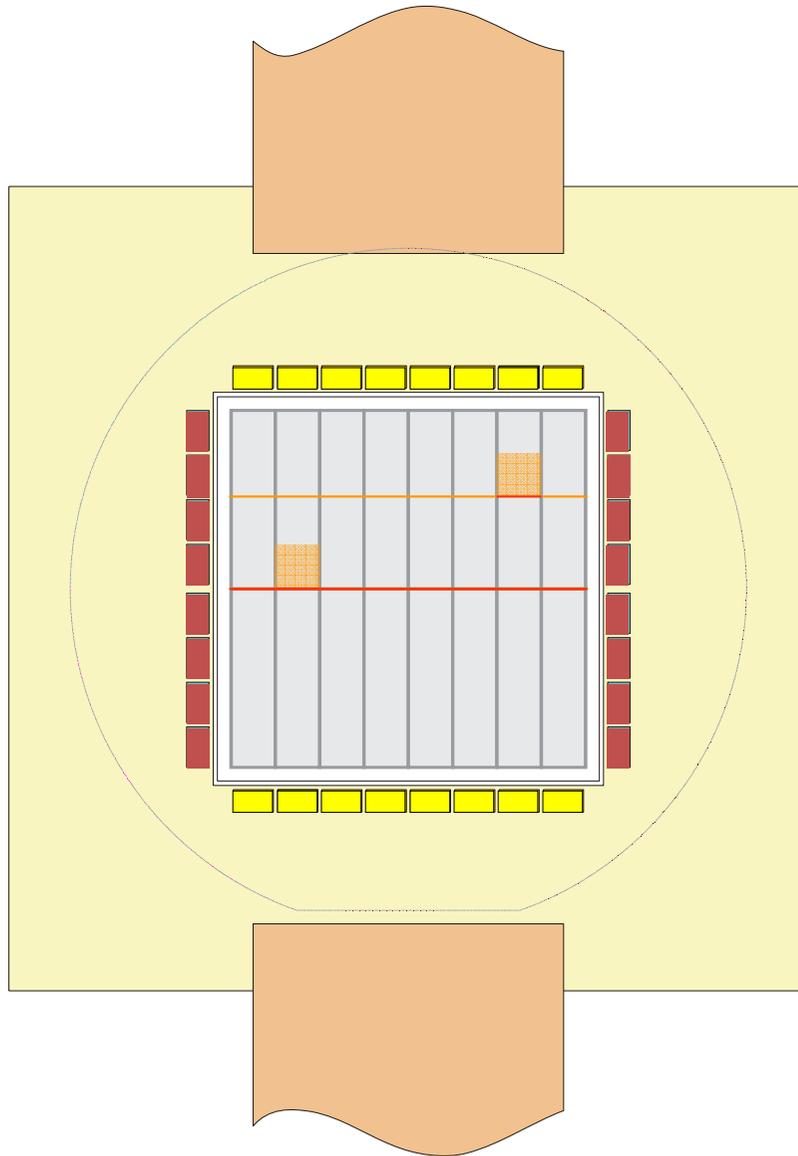
WFI readout concept



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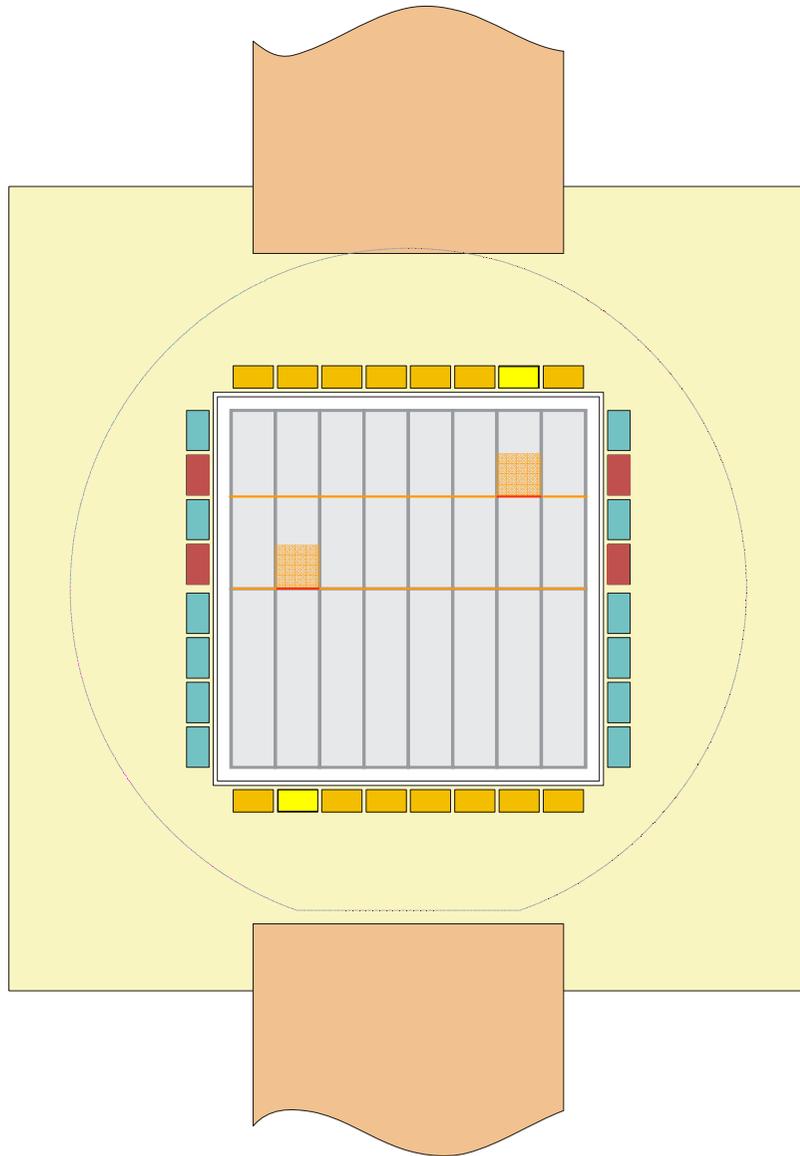
WFI readout concept



■ *Readout modes:*

- **Window mode:** Acquire fully sized window strip (anywhere on FPA) repetitively
- Read rest of frame with reduced framerate
- Different ROIs on arbitrary positions on FPA
- Even different non-overlapping ROIs on same Hemisphere possible (subsequent readout)

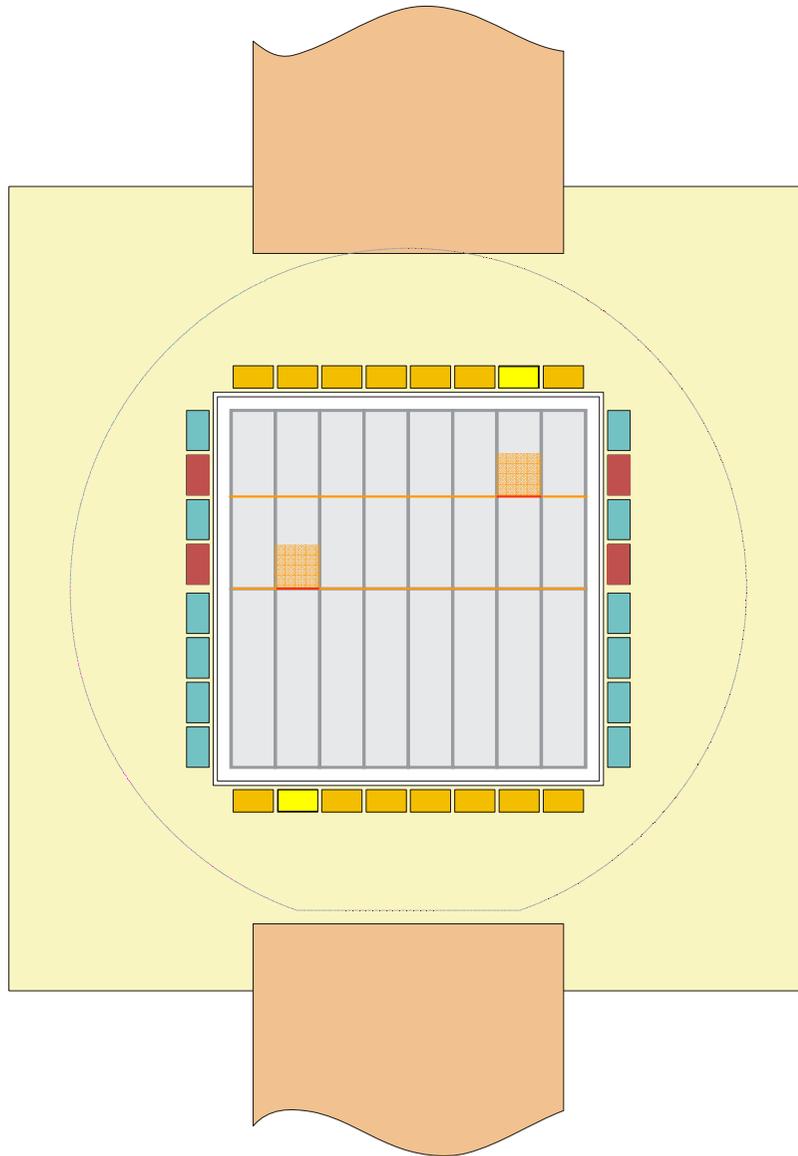
WFI readout concept



- *Readout modes:*

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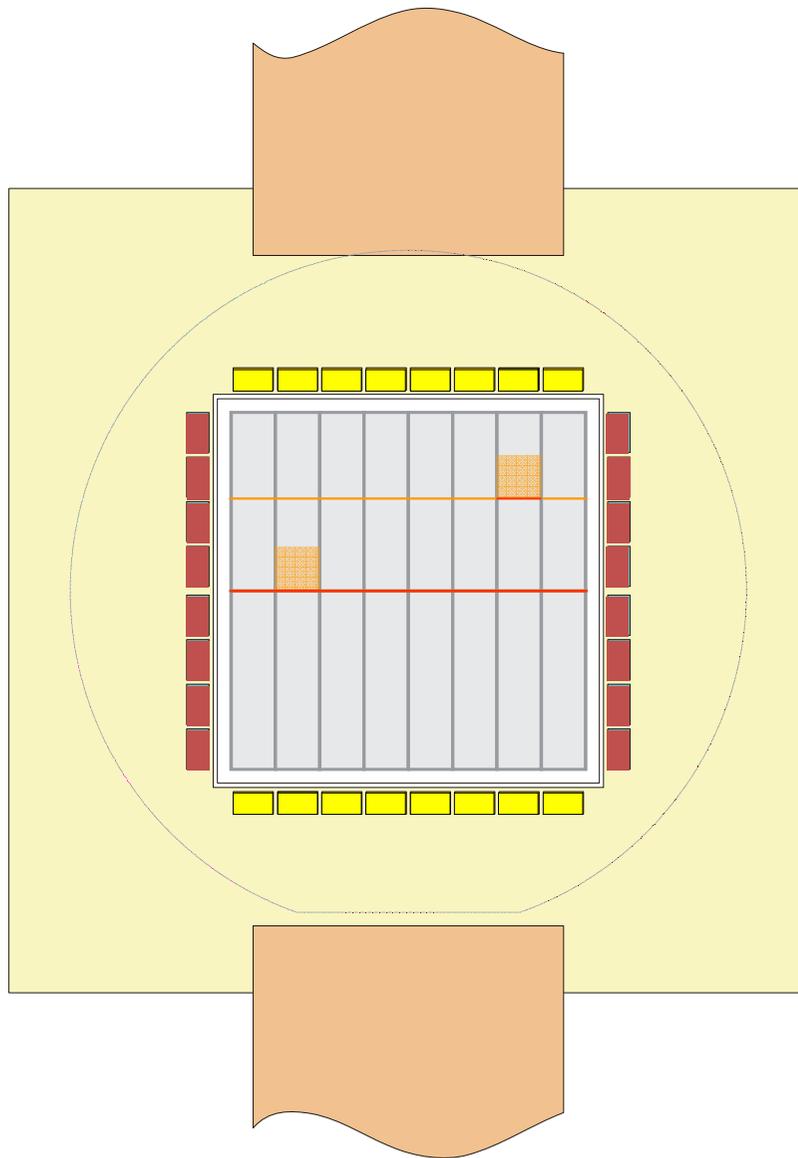
WFI readout concept



■ *Readout modes:*

- **Window mode:** Acquire fully sized window strip (anywhere on FPA) repetitively
- Read rest of frame with reduced framerate
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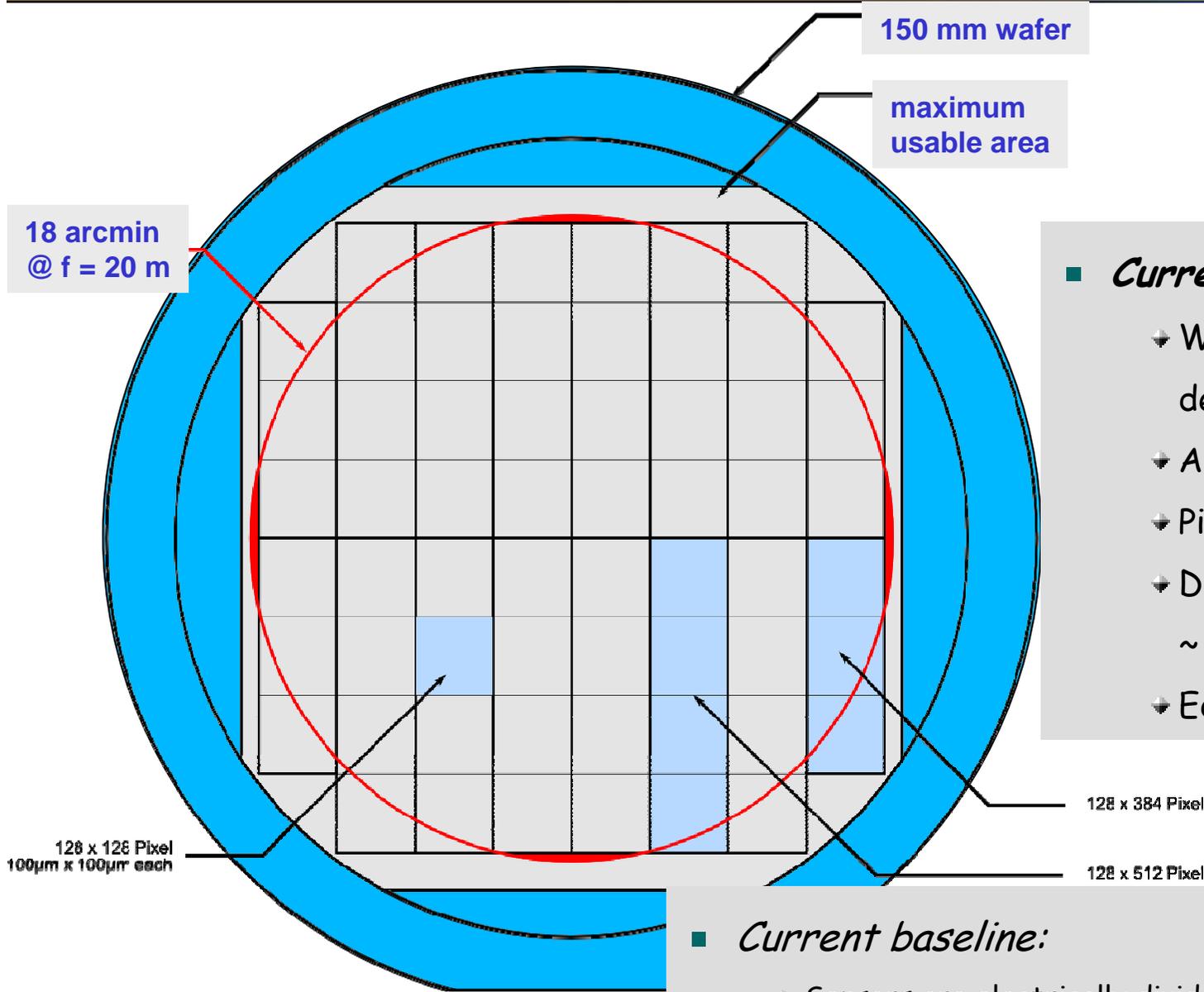
WFI readout concept



- *Readout modes:*

- **Window mode (cont.):** Acquire fully sized window strip (anywhere on FPA) repetitively
- Read rest of frame with reduced framerate
- Different ROIs on arbitrary positions on FPA
- Even different non-overlapping ROIs on same Hemisphere possible (subsequent readout)

Concept



■ **Current baseline:**

- WFI will consist of monolithic device integrated on 150 mm wafer
- Array format: 1024 x 1024
- Pixel size of $\sim 100 \times 100 \mu\text{m}^2$
- Device will occupy an area of about $\sim 102.4 \times 102.4 \text{ mm}^2$
- Edge regions require special care!

■ **Current baseline:**

- Sensors are electrically divided into two hemispheres and 16 sectors
- Division is purely electrical, Processing time $\sim 2 \mu\text{s} / \text{row}$
- Common entrance window homogeneous & without dead regions
- Hemispheres are read out in parallel (factor 2 in parallelization)

Expected and experimentally verified WFI properties

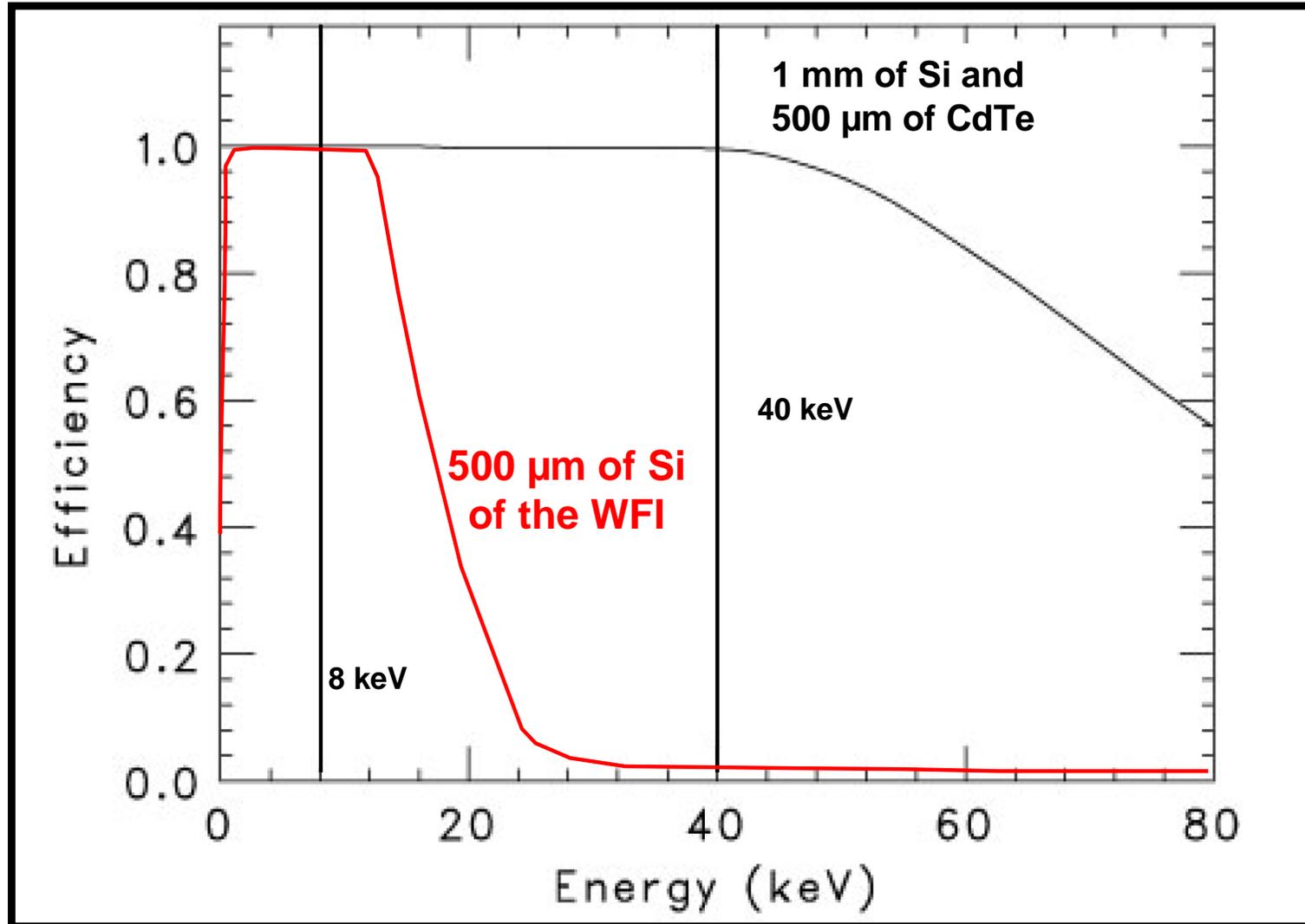
- | | |
|-------------------------------|--|
| 1. Energy bandwidth: | 0.1 keV up to 20 keV |
| 2. Electronic noise: | 3 electrons (rms) |
| 3. Energy resolution: | 130 eV (FWHM) @ 6 keV
45 eV (FWHM) @ 0.2 keV |
| 4. Format: | 1024 x 1024 pixels |
| 5. Pixel (FP) size: | 100 x 100 μm^2 (10 x 10 cm^2) |
| 6. Position resolution: | $\sigma_{x,y} \leq 40 \mu\text{m}$ |
| 7. Readout speed | 2 $\mu\text{s}/\text{pix}$, 2048 pix. in parallel |
| 8. Time resolution: | 1 ms in FF, down to 16 μs in WM |
| 9. Integrated optical filter: | 10^5 optical light reduction |
| 10. Windowing modes: | adjustable according to target |

TRL schedule summary

Topic	2009 1. half	2009 2. half	2010 1. half	2010 2. half	2011 1. half	2011 2. half	2012 1. half	2012 2. half
DePFET sensor fabrication	Yellow	Light Green	Light Green	Light Green	Bright Green	Bright Green	Dark Green	Dark Green
CAMEX ASIC	Light Green	Bright Green	Bright Green	Bright Green	Bright Green	Dark Green	Dark Green	Dark Green
VELA & ASTEROID	Orange	Yellow	Yellow	Light Green	Light Green	Bright Green	Bright Green	Bright Green
Switcher ASIC	Yellow	Light Green	Light Green	Light Green	Bright Green	Bright Green	Bright Green	Dark Green
Camera mechanics And HXI I/F	Orange	Yellow	Yellow	Yellow	Light Green	Light Green	Bright Green	Bright Green
Cooling	Orange	Orange	Yellow	Yellow	Light Green	Light Green	Bright Green	Bright Green
optical blocking Filters	Yellow	Light Green	Light Green	Light Green	Bright Green	Bright Green	Bright Green	Bright Green
Filter sled	Orange	Orange	Yellow	Yellow	Light Green	Light Green	Bright Green	Bright Green
DAQ Electronics	Yellow	Yellow	Yellow	Light Green	Light Green	Light Green	Bright Green	Bright Green

TRL	3	4	5	6	>6
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Quantum efficiency of WFI and HXI

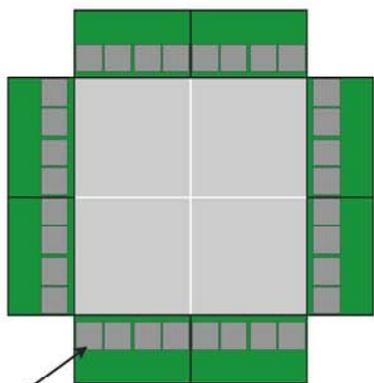


Hard X-ray imaging detector



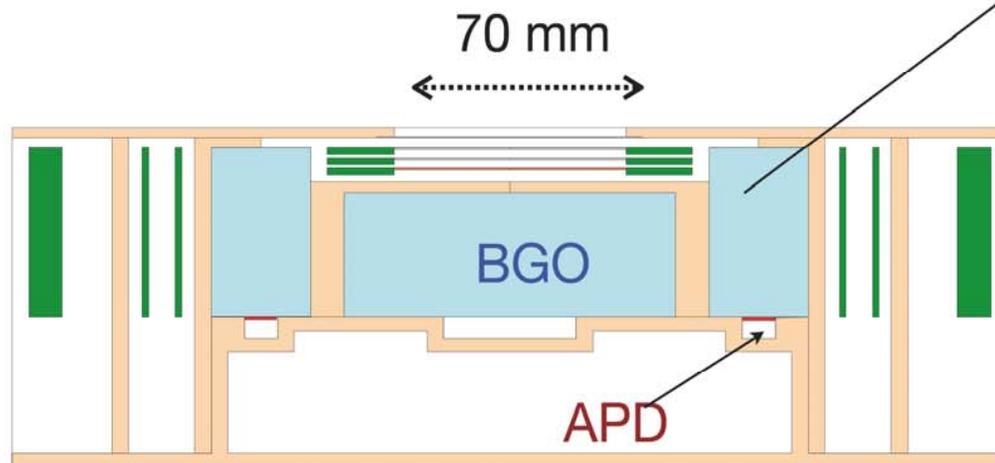
Main detection part consists of 2-layers of Double-Sided Silicon Strip (DSSD) and 1 layer of Double-sided CdTe strip (DS-CdTe). Each layer contains 2x2 devices.

Top-view
(imager part)



Readout ASICs are directly coupled with devices within the same plate.

Cross-section-view



Active shield :
2 cm thick BGO
(default)



Heritage from
Suzaku-HXD

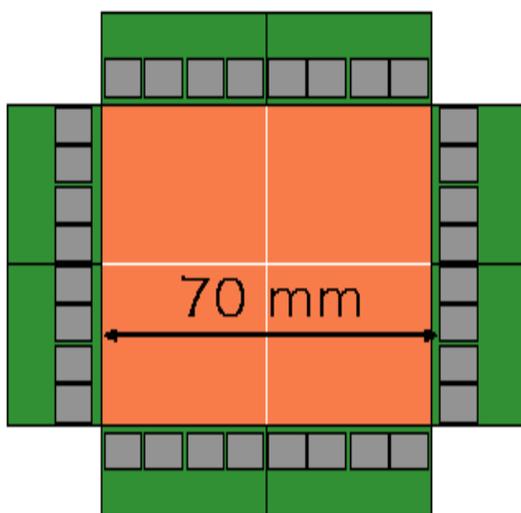
Total Weight : 24 kg
Total Power : ~32W

(including electronics)

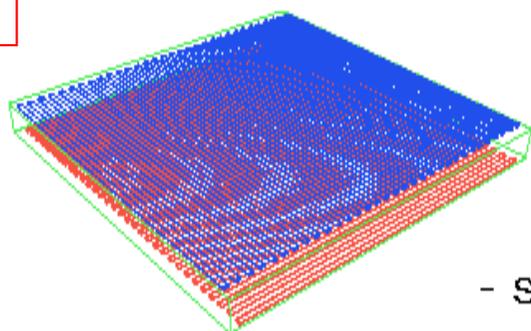
Status of the CdTe imager

Double-sided CdTe strip detector from ACRORAD

IXO
HXI
goal



0.5 (0.75) mmt



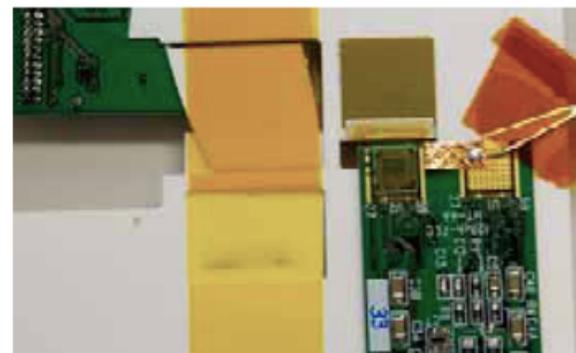
Strip detector

- smaller read-out channels (2n instead of n^2)
- 1-dimensional ASIC
- relatively large leak and C

Laboratory experiment



2.5 cm DS-CdTe



1.3 cm DS-CdTe

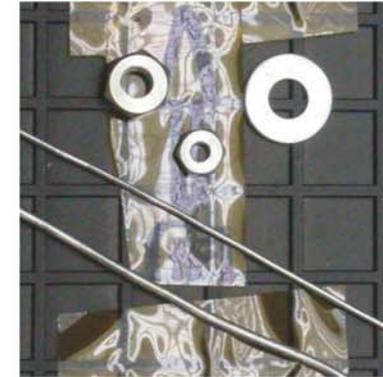
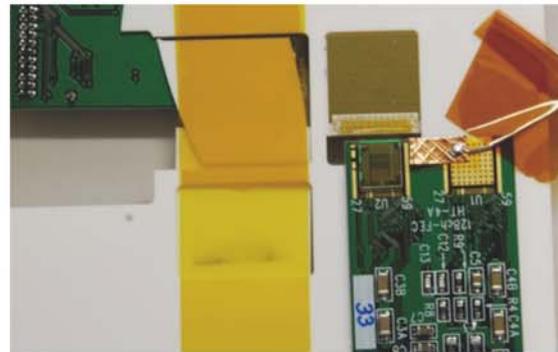
Status of HXI development: DS CdTe

Double-sided strip detector based on CdTe diode devices from ACRORAD

2.5 cm DS-CdTe



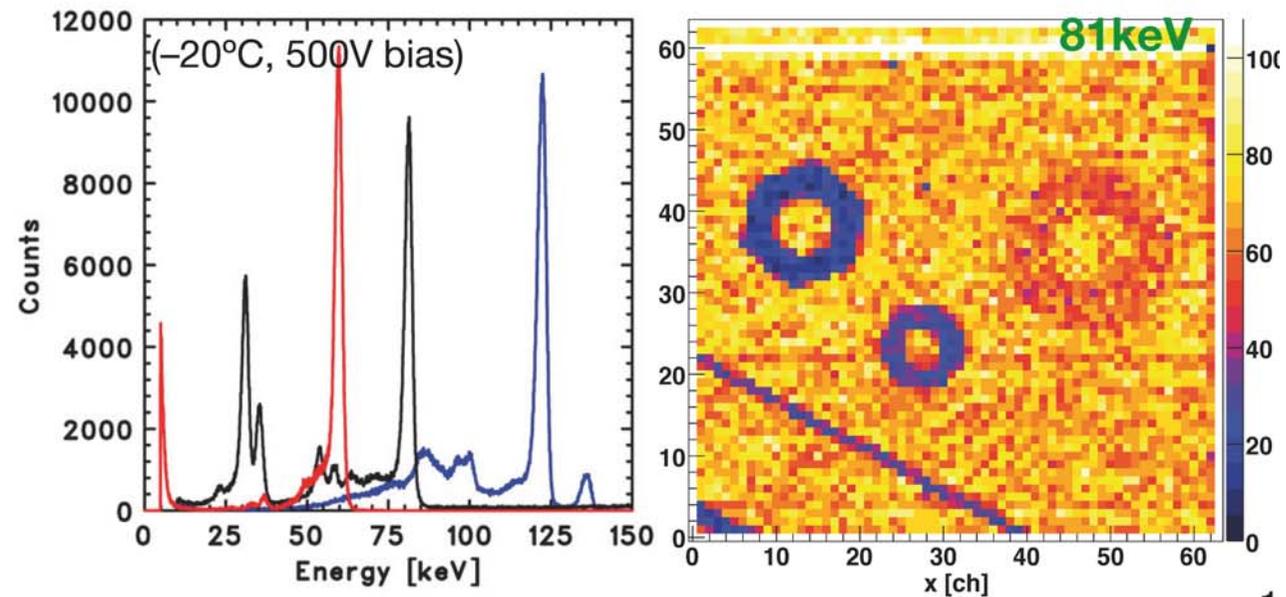
1.3 cm DS-CdTe



Shadow Image with various RIs (^{241}Am , ^{133}Ba , ^{57}Co)

Comparisons of strip with Pixel

- smaller read-out channels (2n instead of n^2)
- 1-dimensional ASIC
- relatively large leak and C



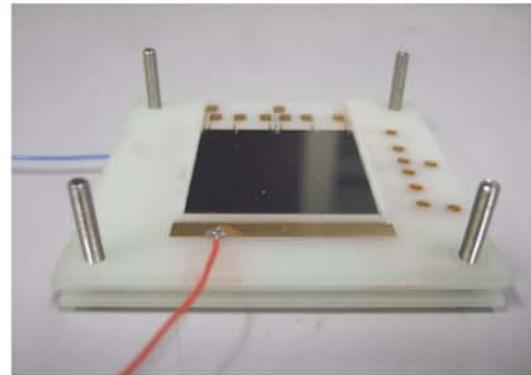
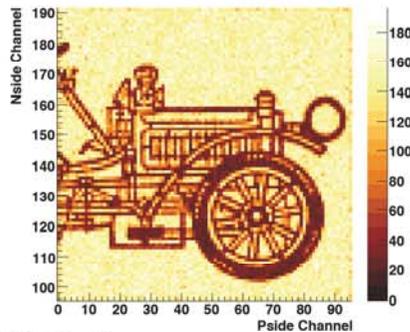
Status of HXI development:

DS SSD

Lower intrinsic background can be achieved than CdTe thanks to its small atomic number. Also act as a “fluorescent X-ray shield” for WFI located between CdTe and WFI. Background reduction with anti-coincidence can be also utilized.

Engineering model

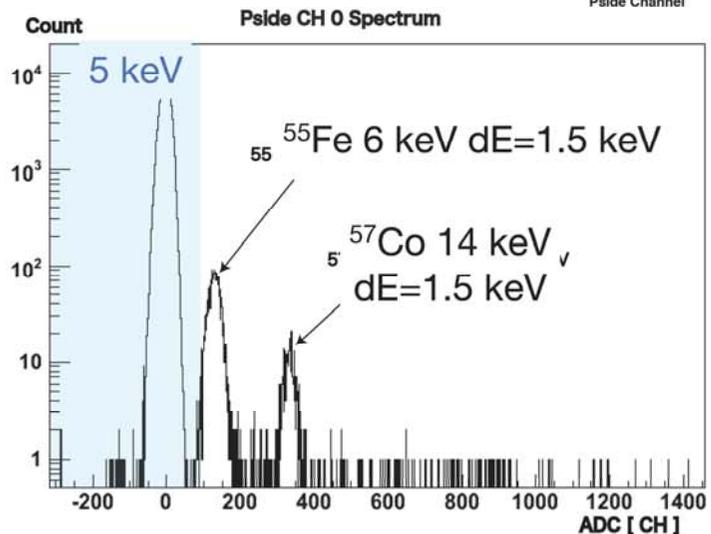
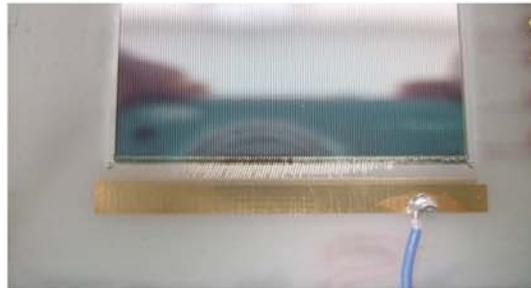
Strip pitch: 400 μ m
Size : 40 x 40 mm²



ASTRO-H flight model

Strip pitch: 250 μ m
Thickness: 500 μ m

Area: 32 x 32 mm
(128 channels for each side)



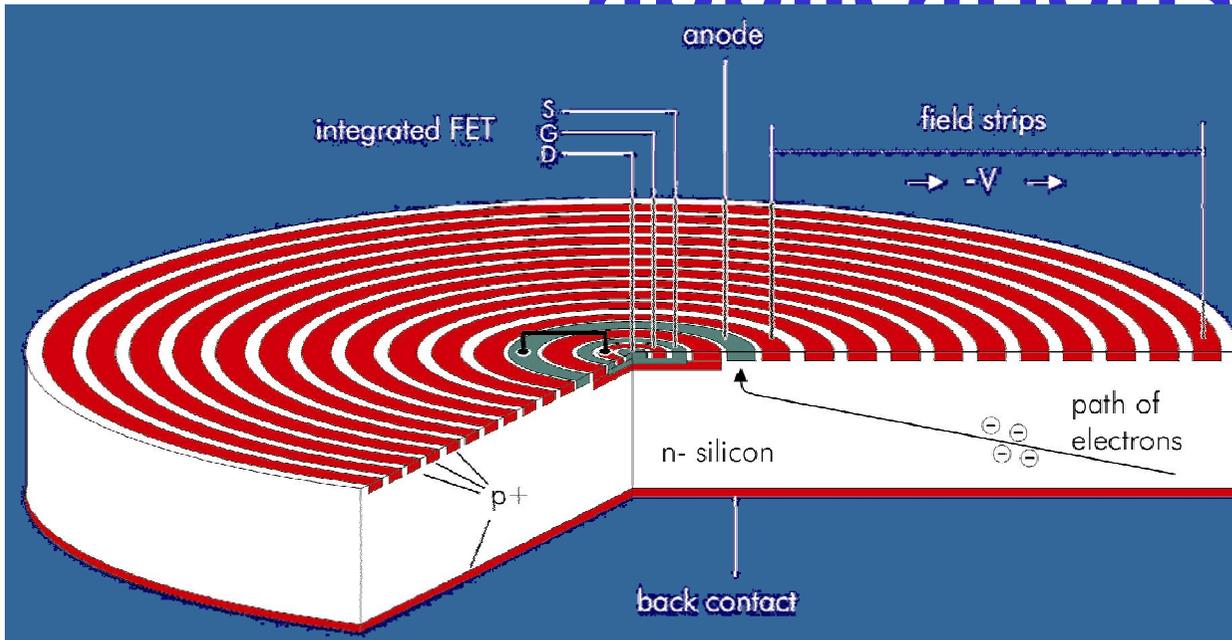
Low noise performance has been verified with a lower energy threshold of ~5 keV.

SDDs for scientific applications

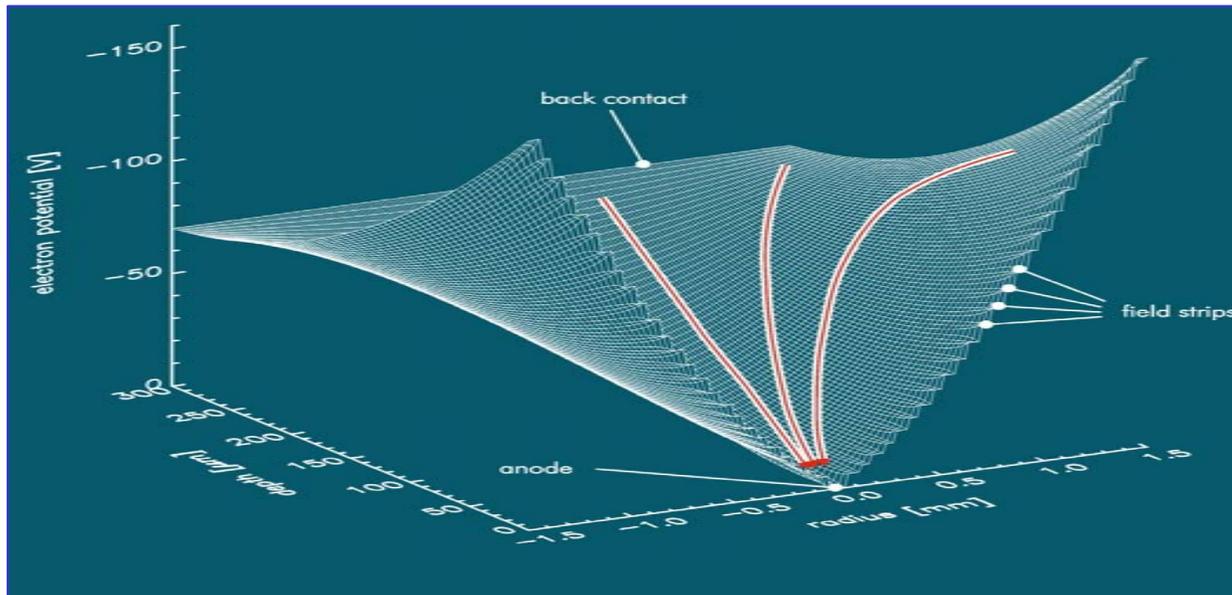
time to drift from device edge to readout node:

$$\Delta t = \Delta x / \mu E = 200 \text{ ns}$$

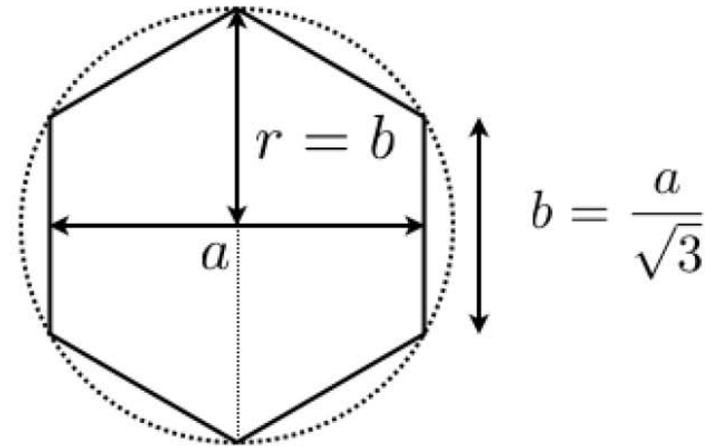
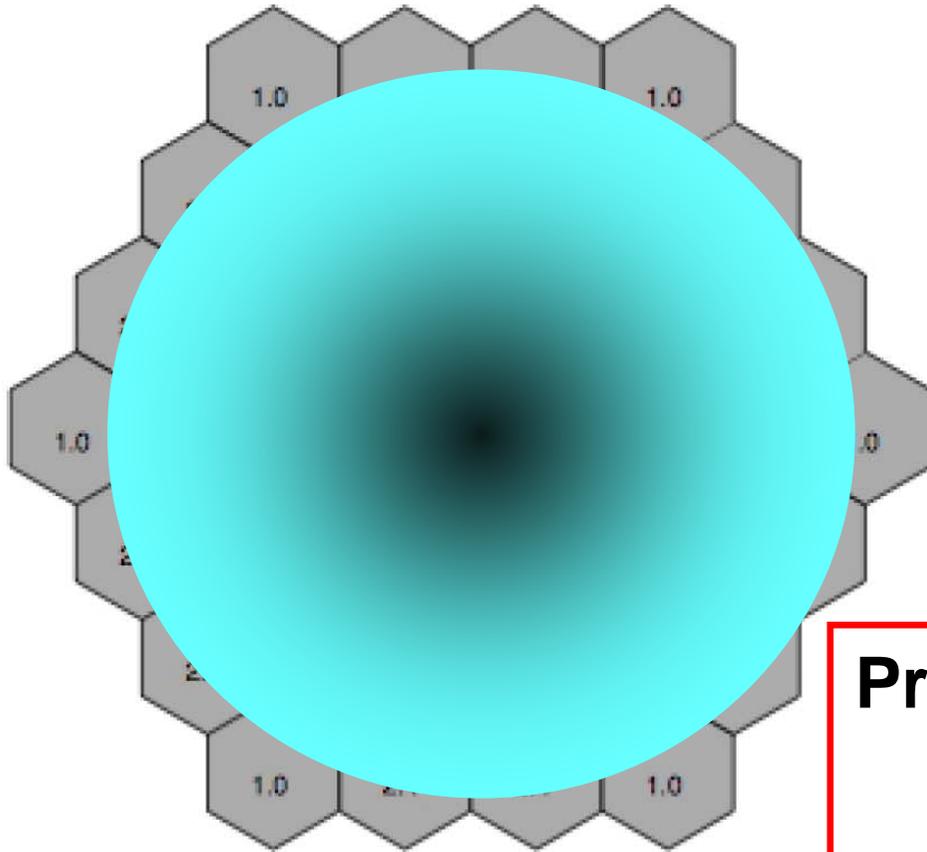
for $\Delta r = 2 \text{ mm}$
i.e. $A = 13 \text{ mm}^2$
and $E = 800 \text{ V/cm}$



SDD with integrated SSJFET



Electrical Potential in a circular SDD



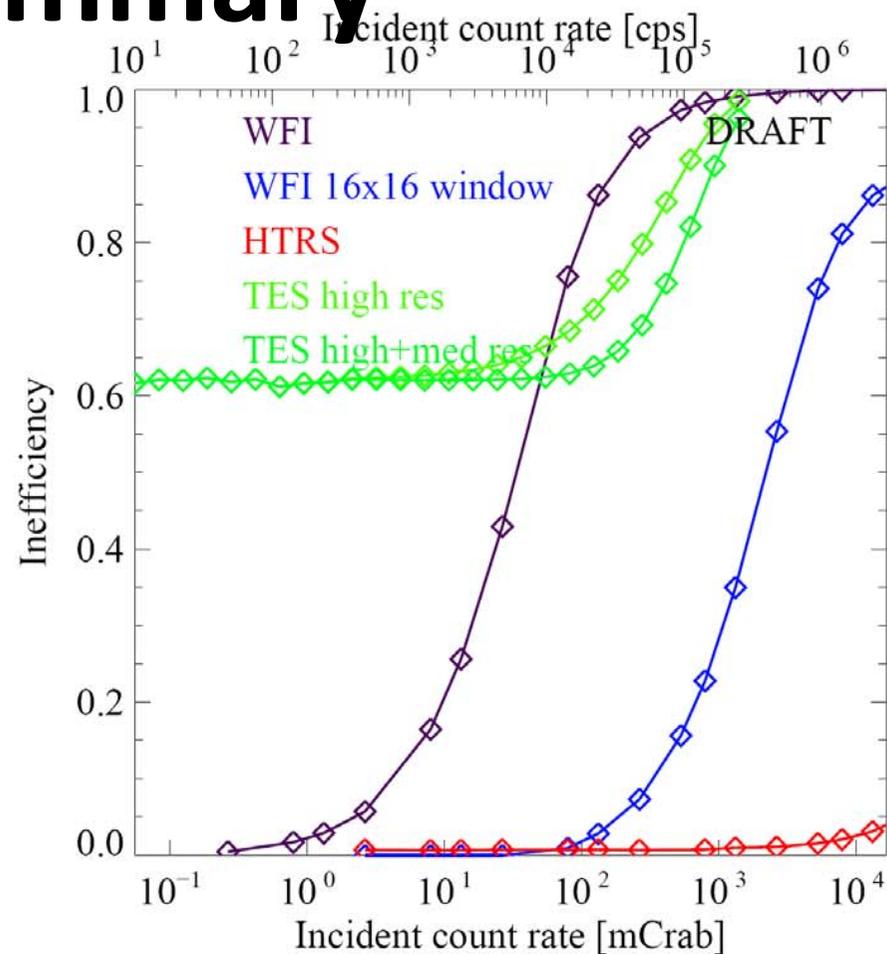
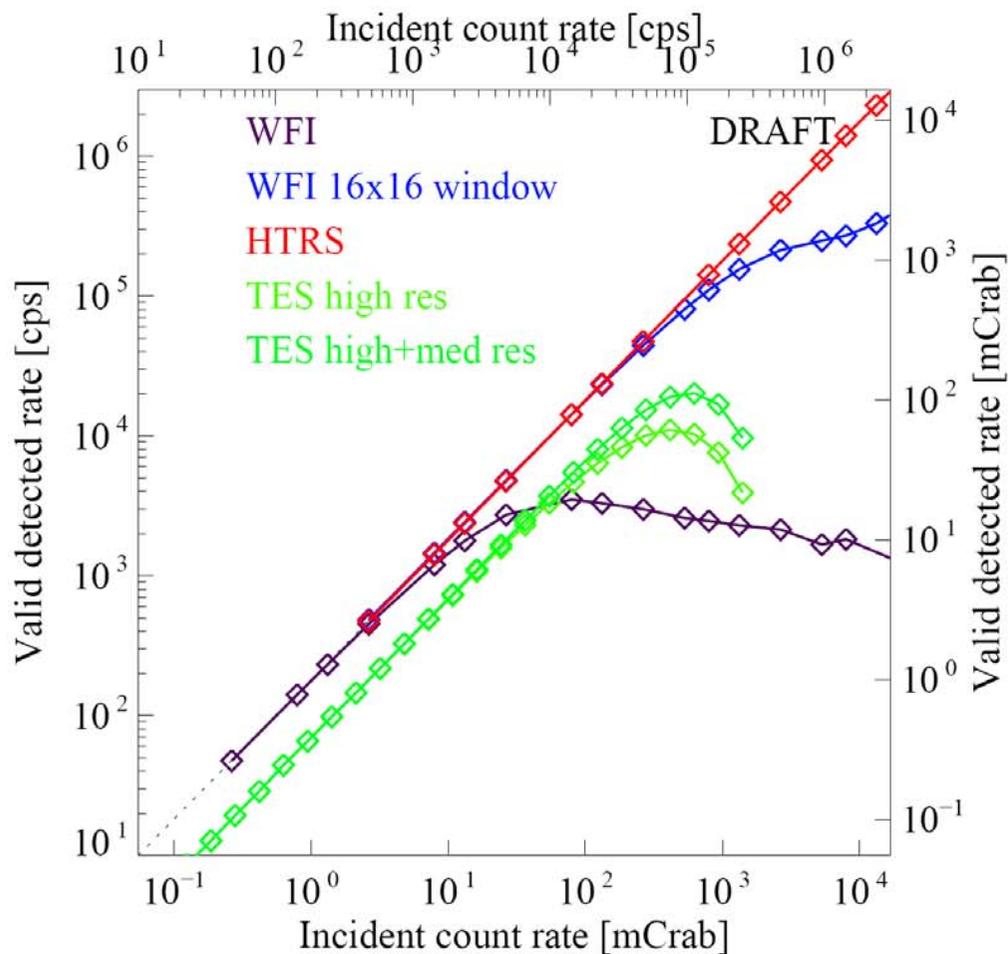
Proposal: array of 37 SDD cells

12 cm out-of-focus
 typ. 10 mm² of
 "adaptive pixel"
 75 ns shaping
 up to 500,000 c
 $\Delta E < 200$ eV (F

500 μ m thick sc

0.1 keV $< E < 2$

Count rate capability: Summary



Calculations performed by Jörn Wilms et al.

with input
from

- **MPI-HLL (MPE and PNSensor)**
- **Universties of Tübingen, Darmstadt and Nürnberg-Erlangen**
- **University of Leicester**
- **Politecnico di Milano**
- **University of Osaka**
- **MIT, CfA and Penn State University**

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